

*International Conference on the Structure and the Interactions of the Photon
including the 19th International Workshop on Photon-Photon Collisions
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Forward neutrons and photons at HERA



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DESY



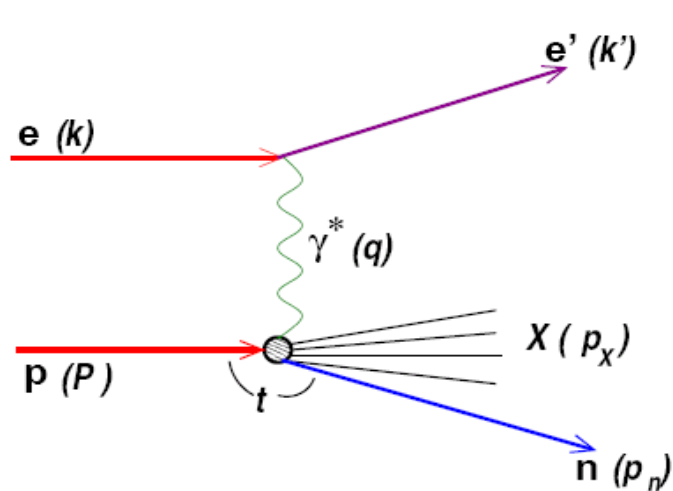
On behalf of the H1 and the ZEUS Collaborations

- **Outline:**
- **Forward neutron p_T and x_L distributions in DIS**
- **Forward neutrons in dijet photoproduction**
- **Forward photon spectra in DIS**

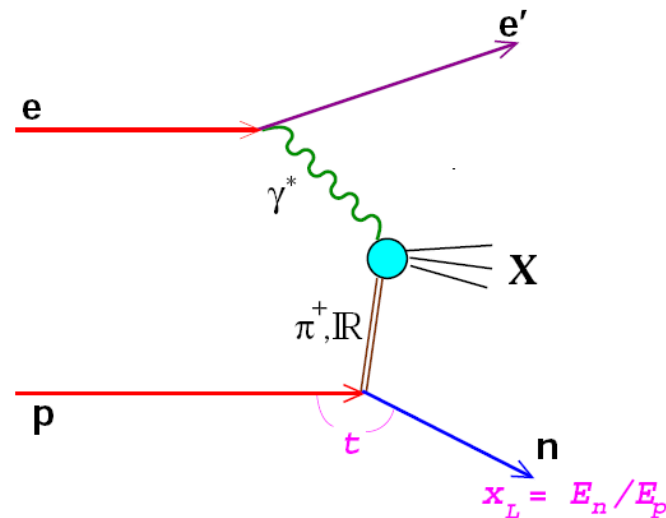
Introduction

Significant fraction of ep scattering events contain a leading neutron (called sometimes forward neutron) in the final state carrying a substantial portion of the energy of the incoming proton:

$e+p \rightarrow e+n+X$. Different production mechanisms are available:

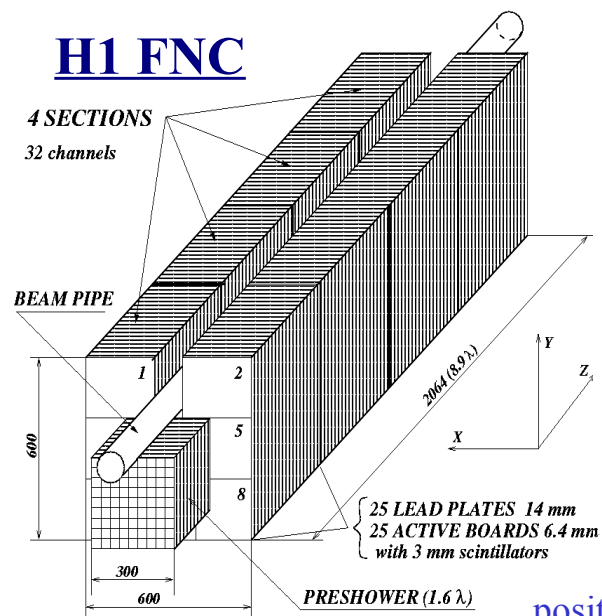
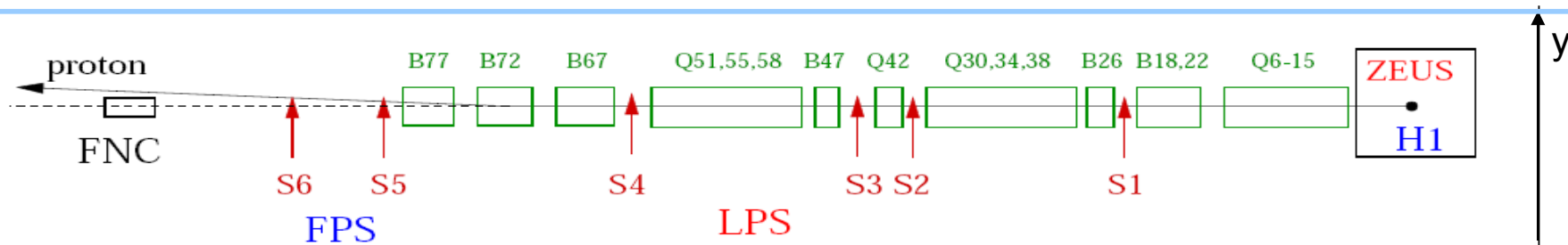


Fragmentation of proton



Exchange of virtual particle: π^+, ρ^+, a_2^+

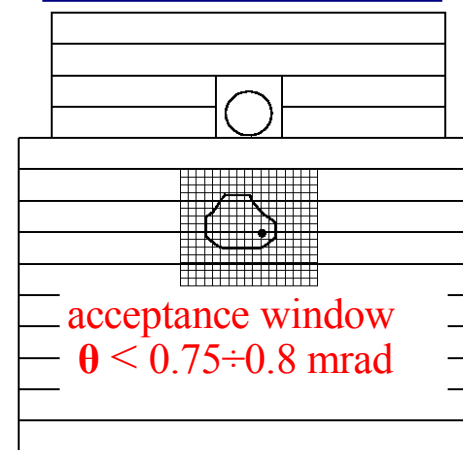
Detectors of Forward Neutrons and Photons



$$\sigma_E/E \approx 0.63/\sqrt{E} \oplus 2\%$$

position resolution 2-3 mm

ZEUS FNC+FNT

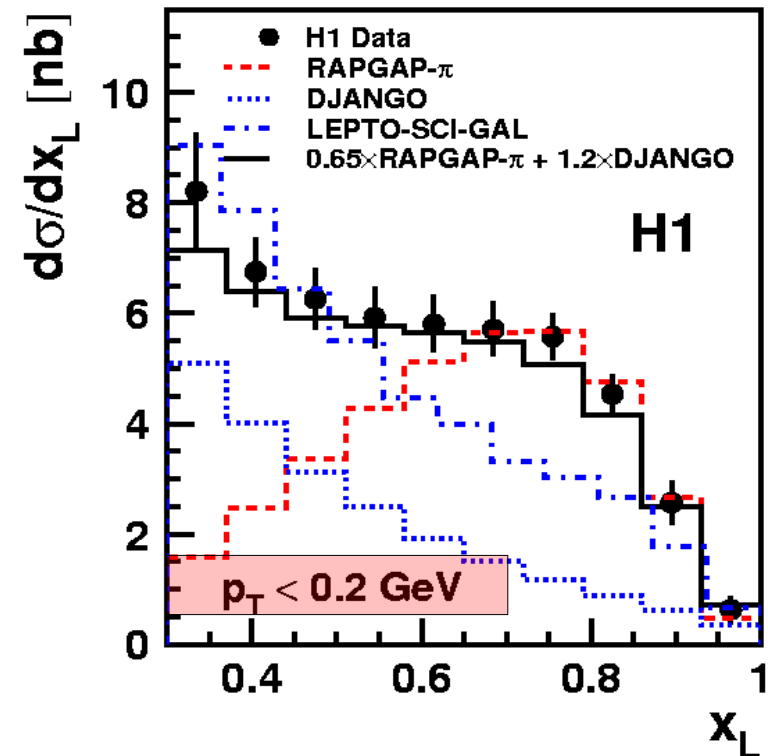
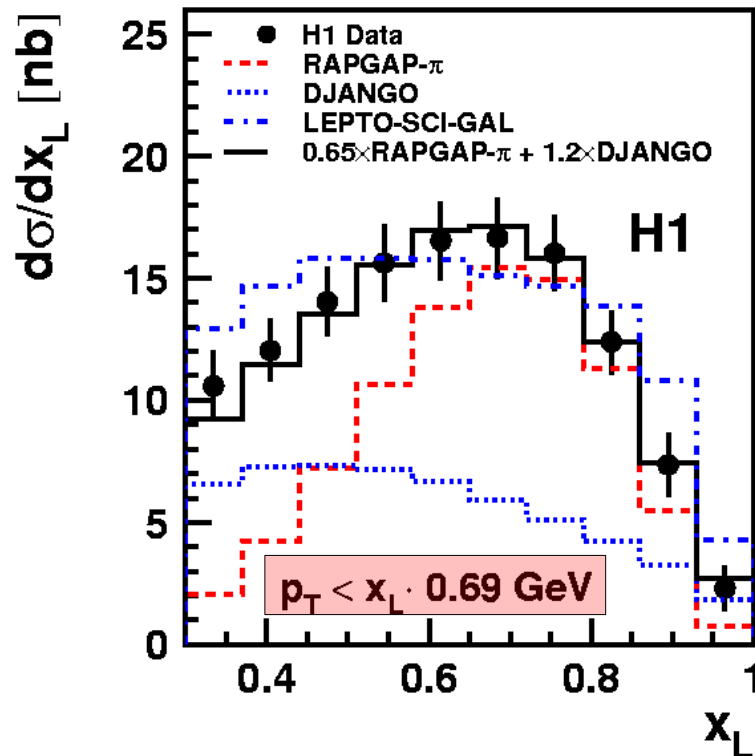


14 towers, 17x15 grid
of the FNT hodoscopes,
 $\sigma_E/E \approx 0.7/\sqrt{E}$

- Segmentation in depth (Z-axis) allows reliable discrimination between electromagnetic and hadronic showers, that is between photons and neutrons.
- Acceptance limited by beam apertures (defined by beam line elements).
- p_T resolution is dominated by p_T spread of proton beam (50-100 MeV).

$$e+p \rightarrow e+n+X$$

(Eur.Phys.J.C68:381-399,2010)



Data can be described by a combination of standard fragmentation (DJANGO-CDM) and pion-exchange (RAPGAP- π)

$$e+p \rightarrow e+n+X$$

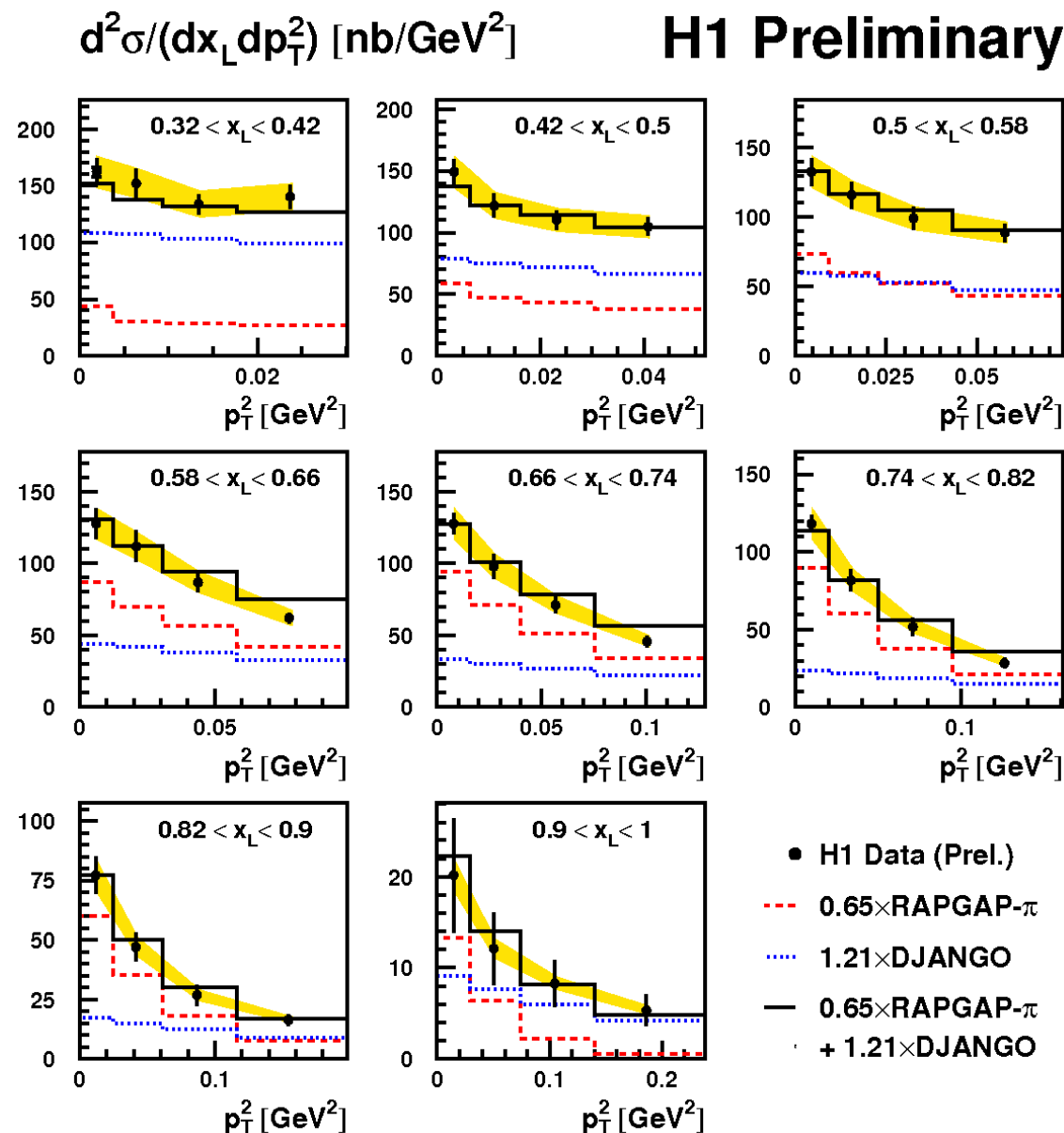
Kinematic range: $6 < Q^2 < 100 \text{ GeV}^2$,
 $0.05 < y < 0.6$, 2006-2007 data,
 $\text{Lumi}=122\text{pb}^{-1}$

Now we extend the measurement
differentially in transverse
momentum of neutron p_T

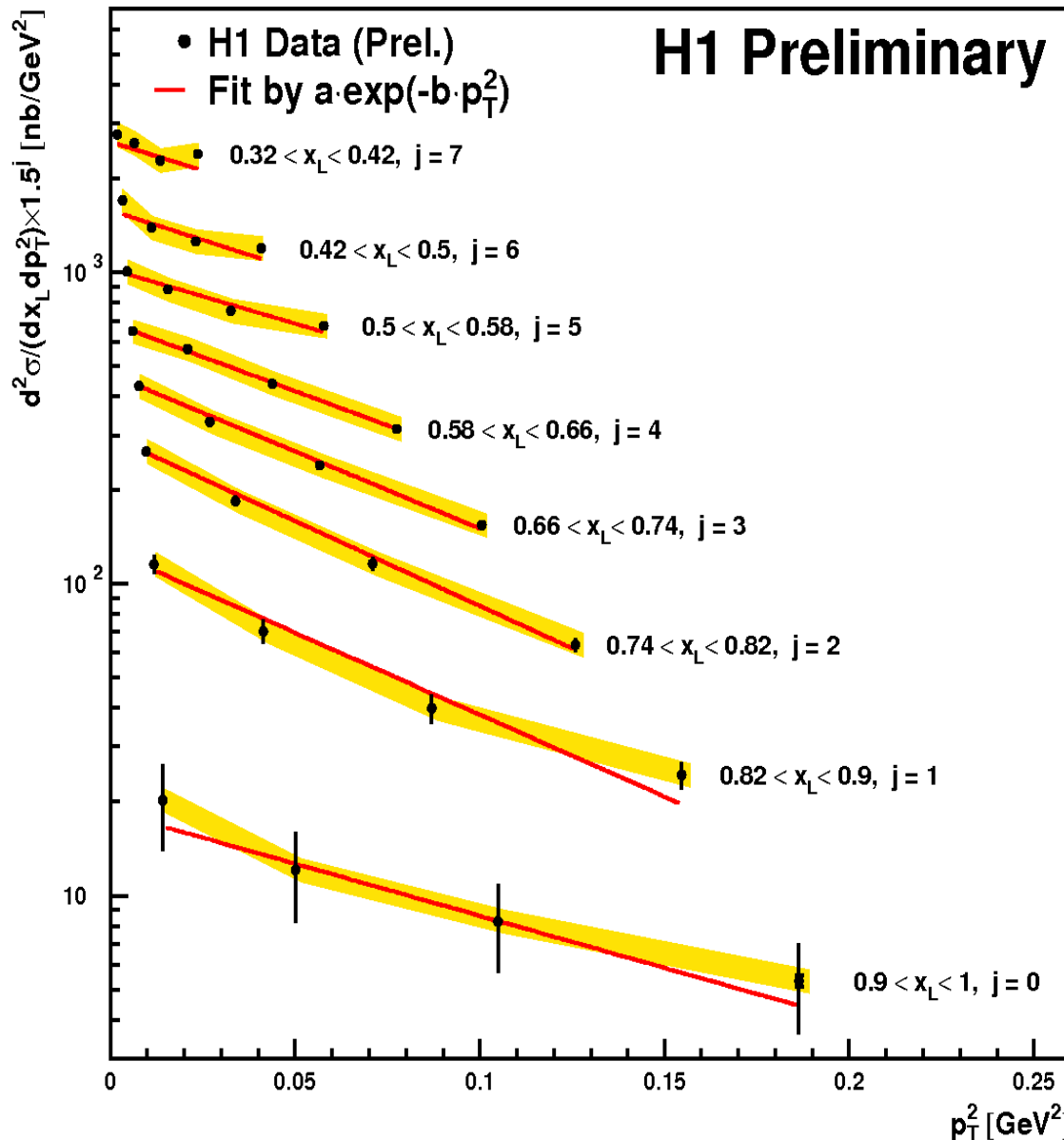
MC weight factors of 0.65 and
1.21 for RAPGAP and DJANGO
correspondingly taken from the
fit to x_L distribution also well
describe p_T^2 distributions

p_T^2 slopes are different for
standard fragmentation and
pion-exchange, constant vs x_L
for DJANGO and increasing
with x_L for RAPGAP

H1 Preliminary



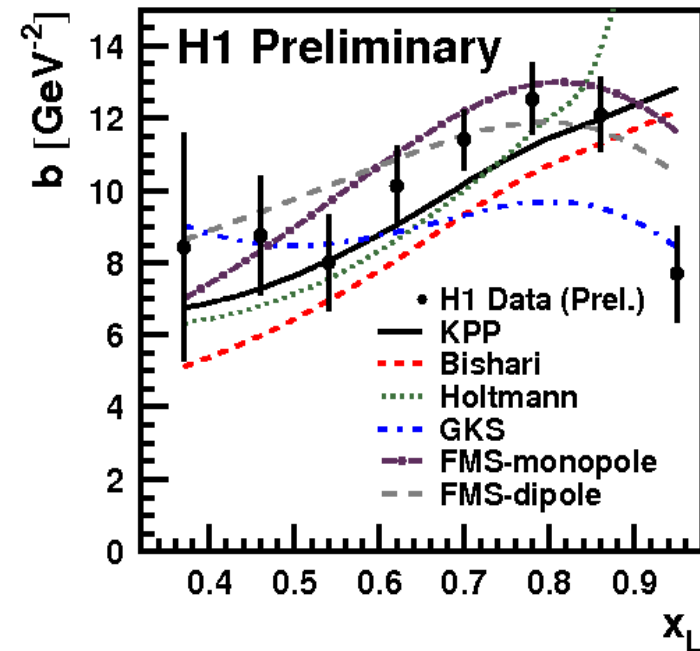
$$e+p \rightarrow e+n+X$$



Fit by: $\frac{1}{\sigma_{inc}} \frac{d\sigma_{LN}}{dp_T^2 dx_L} = a(x_L) \cdot e^{-b(x_L) p_T^2}$

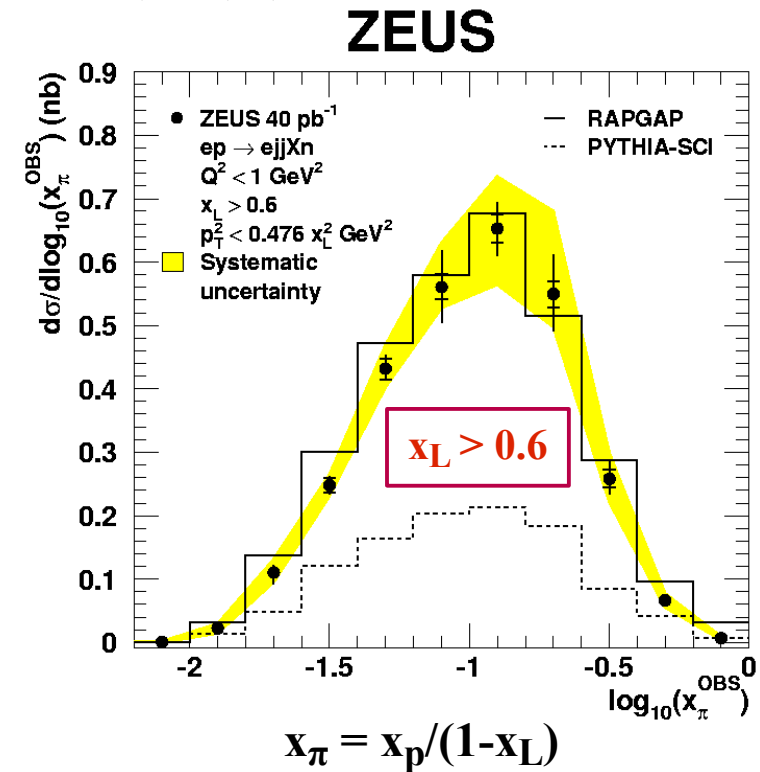
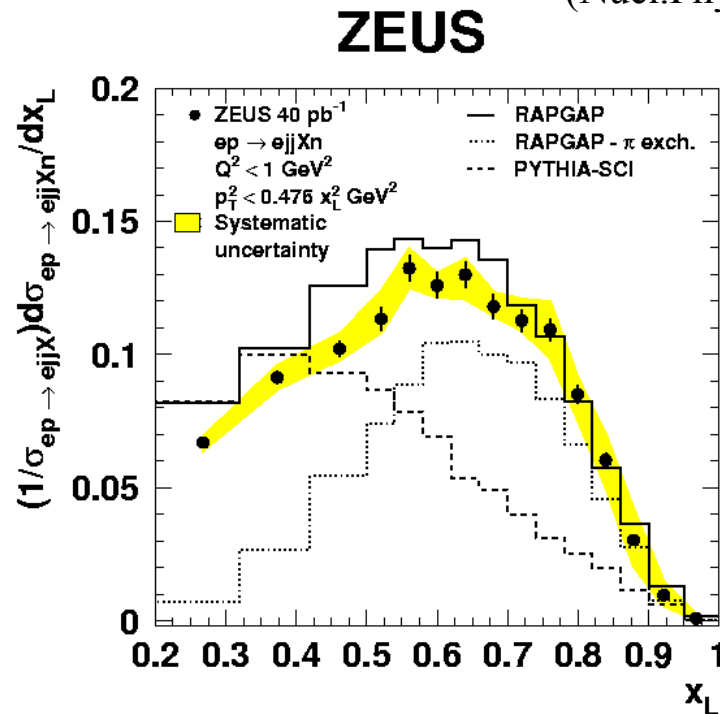
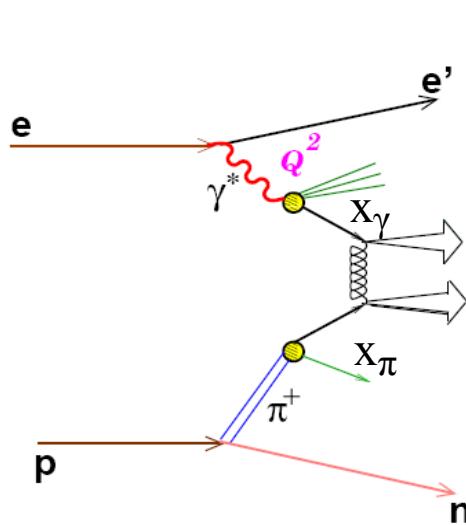
p_T^2 slopes change with increasing p_T^2

Fit the distributions by a single exponent and compare slopes with different parameterizations of pion flux



$$e+p \rightarrow e+n+j+j+X$$

(Nucl.Phys.B827 (2010) 1)



40 pb⁻¹, $Q^2 < 1 \text{ GeV}^2$, $p_T^2 < 0.475 x_L^2 \text{ GeV}^2$, $x_L > 0.2$,
 $130 < W < 280 \text{ GeV}$, $E_T^{\text{jet1}} > 7.5 \text{ GeV}$, $E_T^{\text{jet2}} > 6.5 \text{ GeV}$, $-1.5 < \eta_{\text{jet1,2}} < 2.5$

In photoproduction ($Q^2 \sim 0$) hard scale provided by jets with high p_T^{jet}

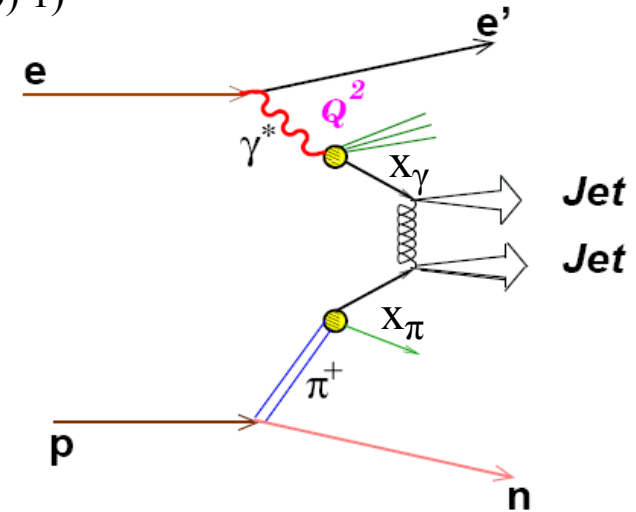
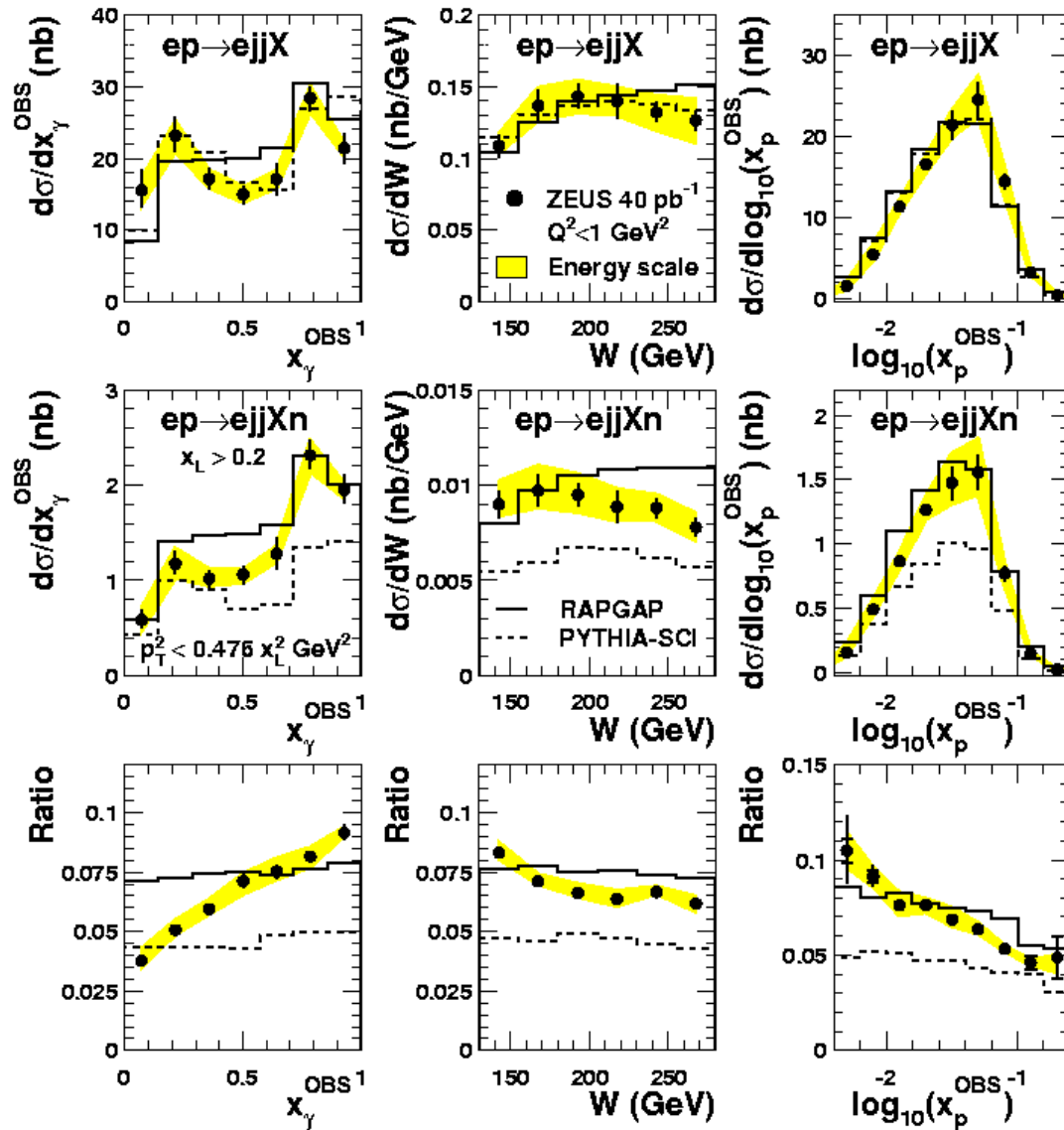
RAPGAP π -exchange and PYTHIA-SCI describe data poor

Pion-exchange is dominating mechanism at high x_L

Full RAPGAP (pion-exchange + inclusive γp) gives good description of data

(Nucl.Phys.B827 (2010) 1)

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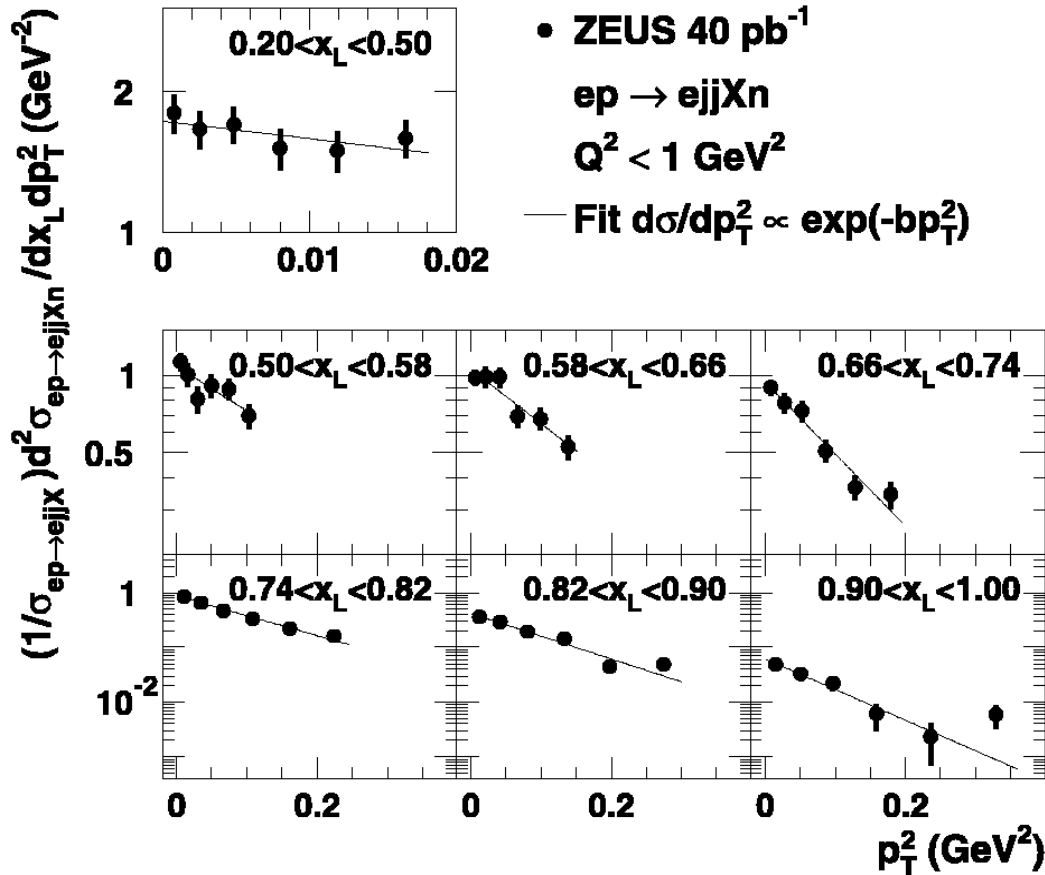
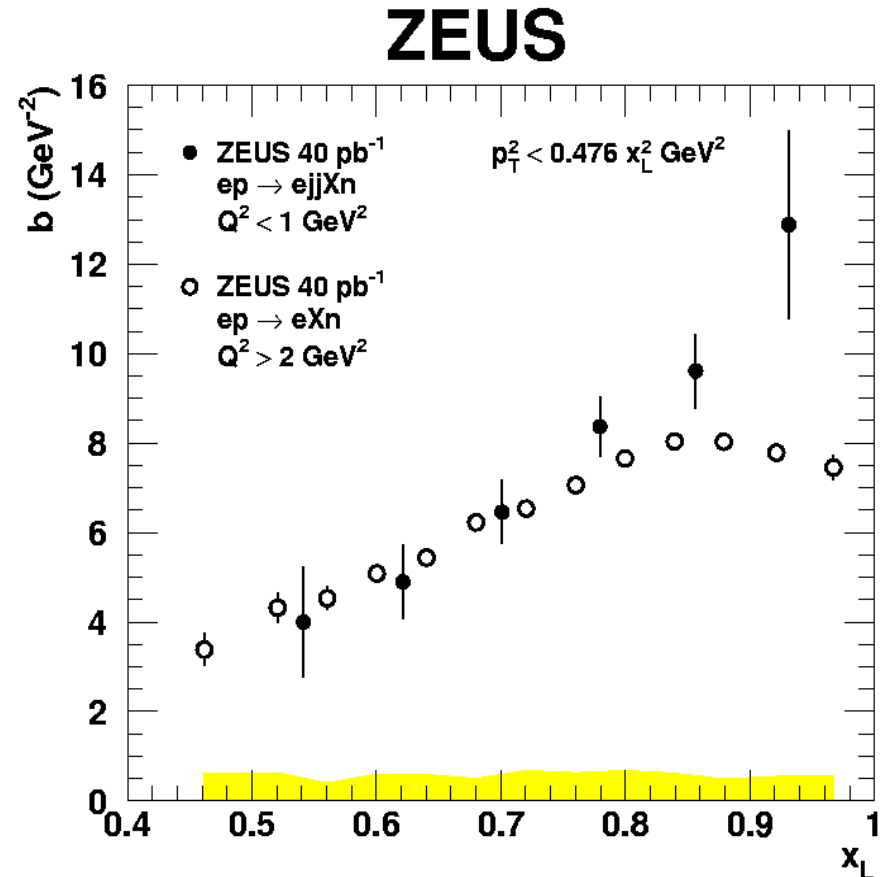


W – total energy of γp system

$$x_\gamma = \sum_{\text{jets}} (E - p_z) / (2yE_e)$$

$$x_p = \sum_{\text{jets}} (E + p_z) / (2E_p)$$

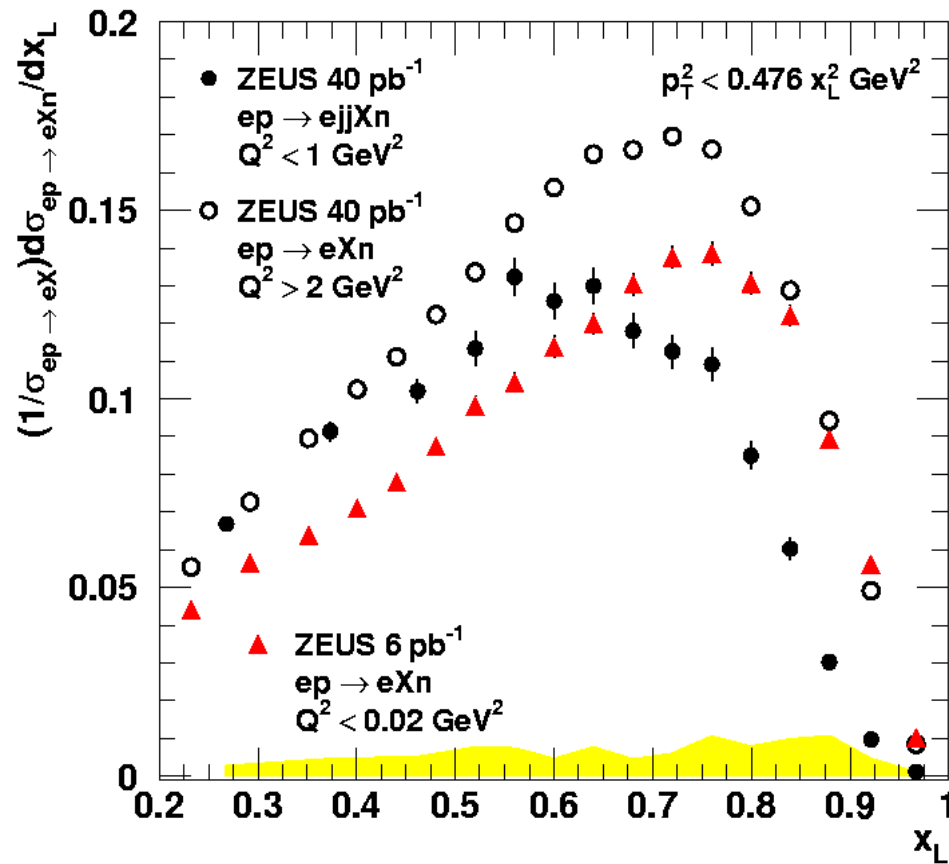
- strong dependence of ratio of x_γ distributions for data, flat in MC
- resolved photon is suppressed in events with neutron.

ZEUS**slopes - $b(x_L)$** Well described by exponential fall-off in p_T^2 

similar b -values in DIS and γp +dijet,
 slightly different at high x_L
 \Rightarrow same production mechanism

(Nucl.Phys.B827 (2010) 1)

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- LN in DIS
- ▲ LN in γp
- LN + dijet in γp

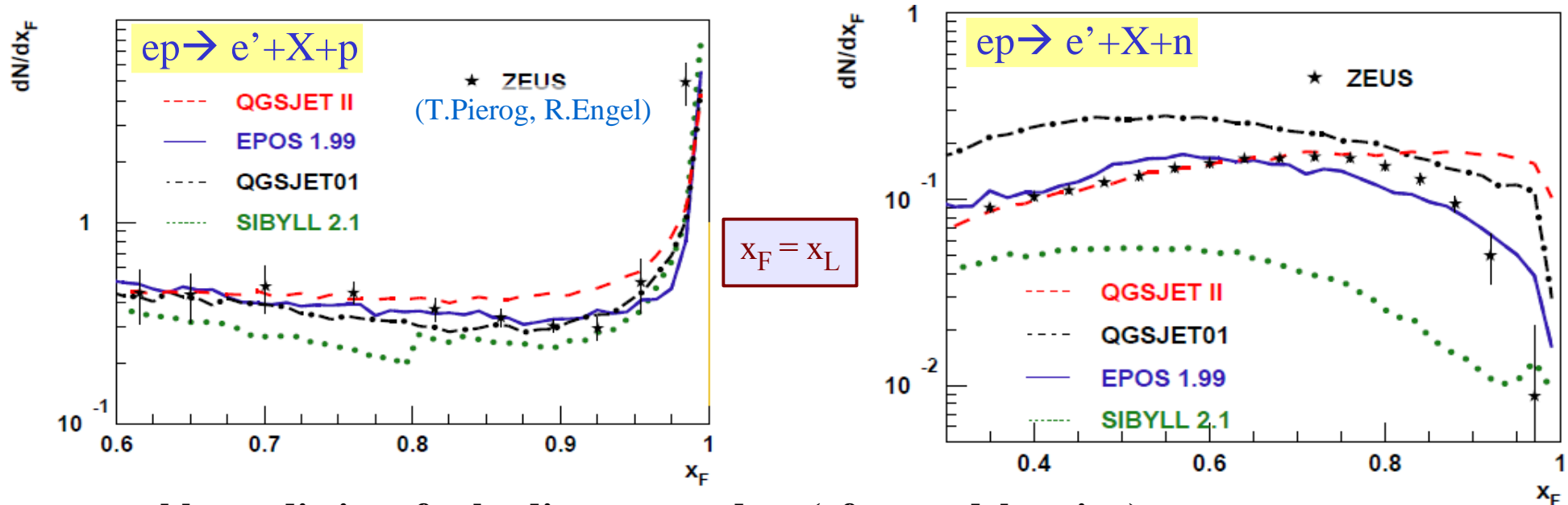
LN production in photoproduction is suppressed vs DIS at lower x_L ($x_L < 0.85$)
=> consistent with neutron absorption through rescattering models (more absorption in γp than in DIS due to larger transverse size of real photon)

Suppression is not so prominent in dijet photoproduction (hard scale provided by high E_T^{jets})

Suppression of dijet photoproduction rate at higher x_L is due to phase space limitation: dijets in the final state leave little room for energetic neutrons

Forward particles production

- Forward particles sensitive to proton fragmentation.
- Forward particles at HERA can contribute to the understanding of high energy cosmic rays



- ♦ reasonable predictions for leading proton data (after model tuning)
- ♦ none of models describe leading neutron data well
- ♦ *What about π^0 ? => measure photons*

Data compared to

- CDM and LEPTO MC
- Hadronic interaction models used for analysis of cosmic rays
EPOS, SIBYLL, QGSJET II, QGSJET 01

(EPOS 1.9: (Pierog, Werner), QGSJET 01 and II: (Kalmykov, Ostapchenko), SIBYLL 2.1: (Engel, Fletcher, Gaisser, Lipari, Stanev))

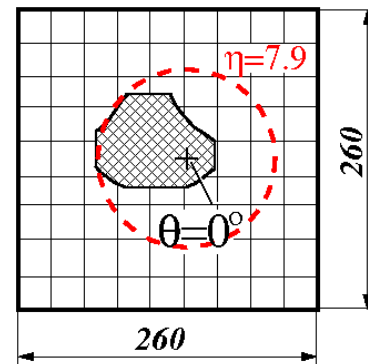
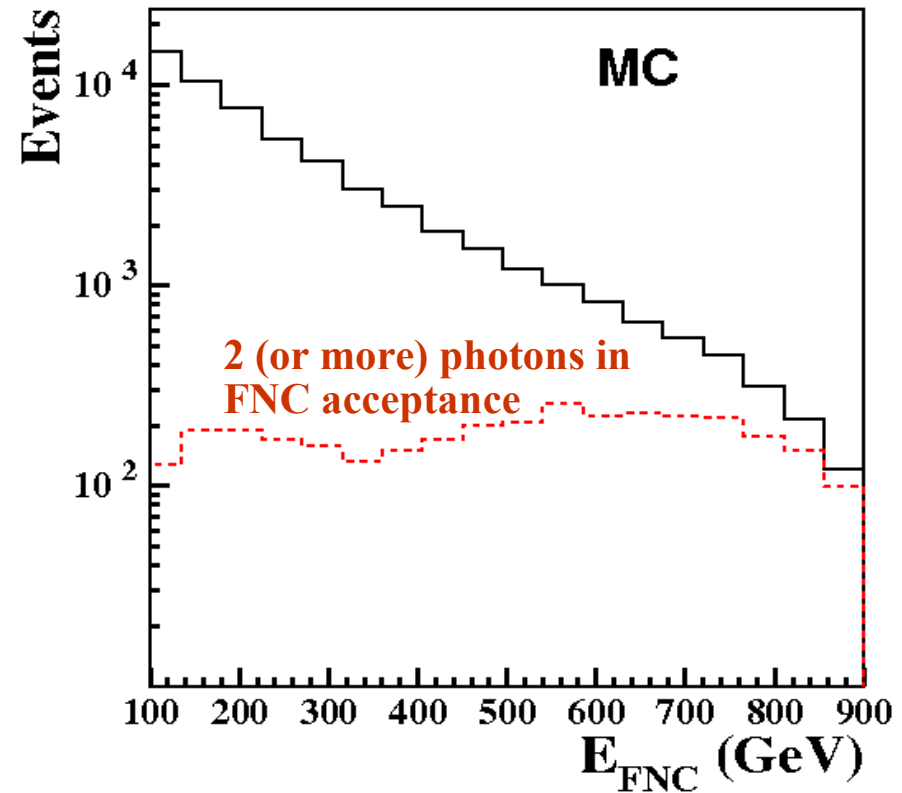
What are our photon candidates ?

At high x_L , many FNC clusters are from more than one photon!
So measurement represents the sum of photons inside the angular range defined by the FNC geometrical acceptance ($\eta > 7.9$).

But at lower x_L we can assume that to a good approximation we measure one photon.

provide cross sections:

- x_L and p_T of most energetic photon in a range $\eta > 7.9$ for $x_L < 0.7$
- x_L of sum of photons in a range $\eta > 7.9$



Forward photons:

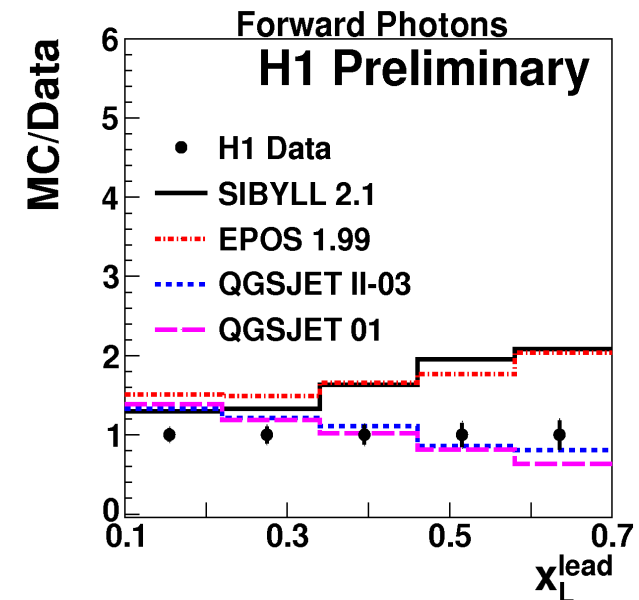
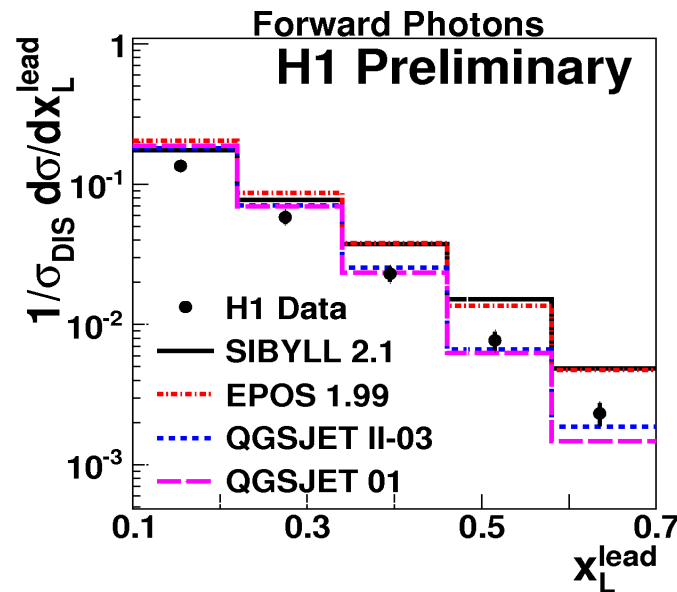
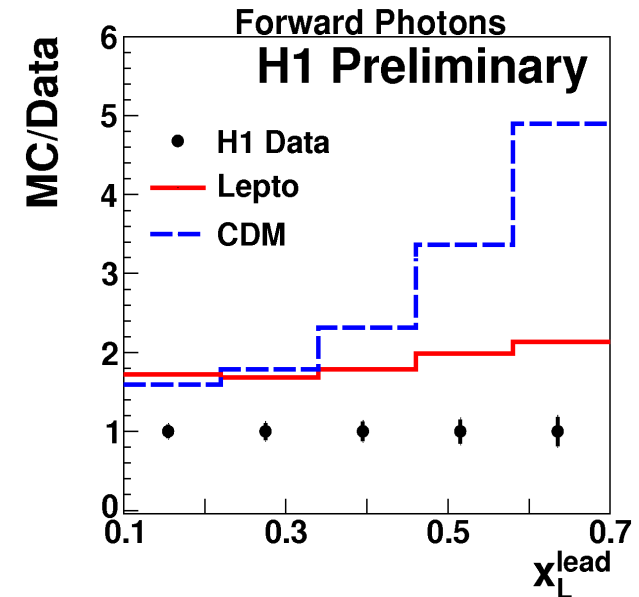
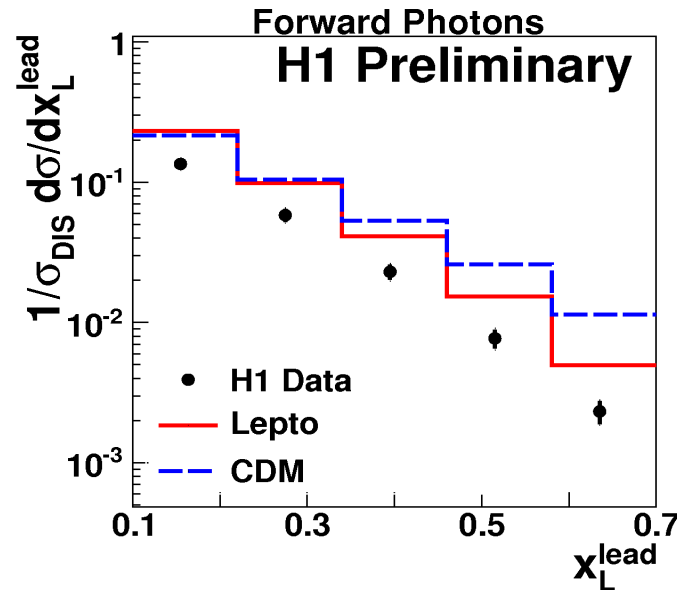
Kinematic range: $6 < Q^2 < 100 \text{ GeV}^2$,
 $0.05 < y < 0.6$, 2006-2007 data,
Lumi= 126 pb^{-1} . σ_{DIS} is inclusive cross
section in the same kinematic range

**Photon rate in all tested
Monte Carlo models is
significantly higher than
in data.**

**LEPTO model describes
the shape reasonably well.
CDM to data discrepancy
larger at higher x_L**

**QGSJET models describe
data well except at low x_L**

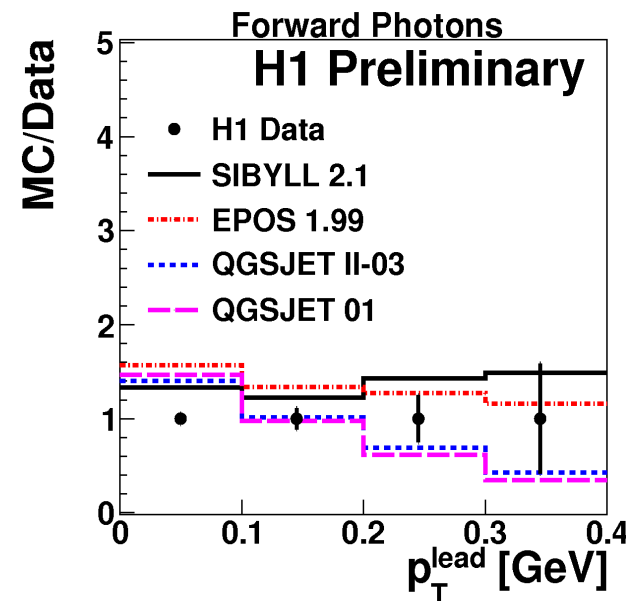
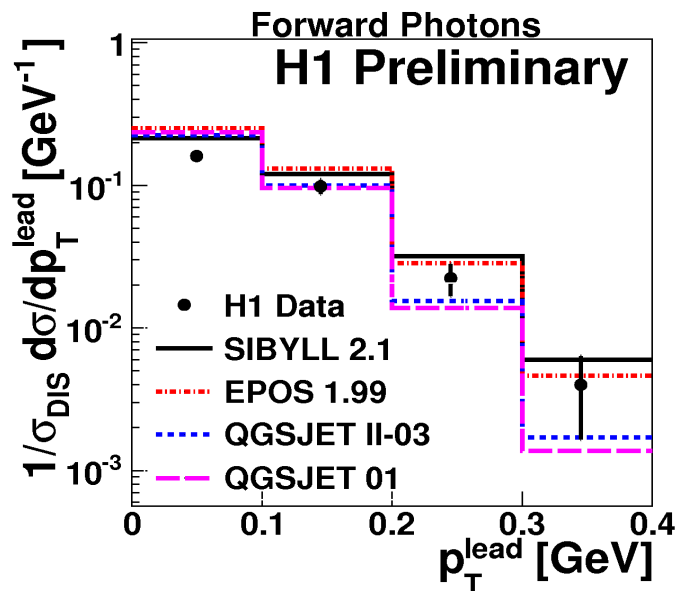
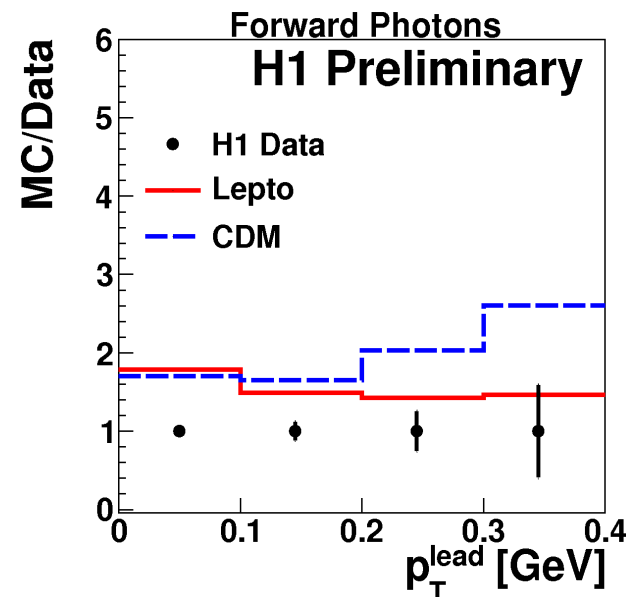
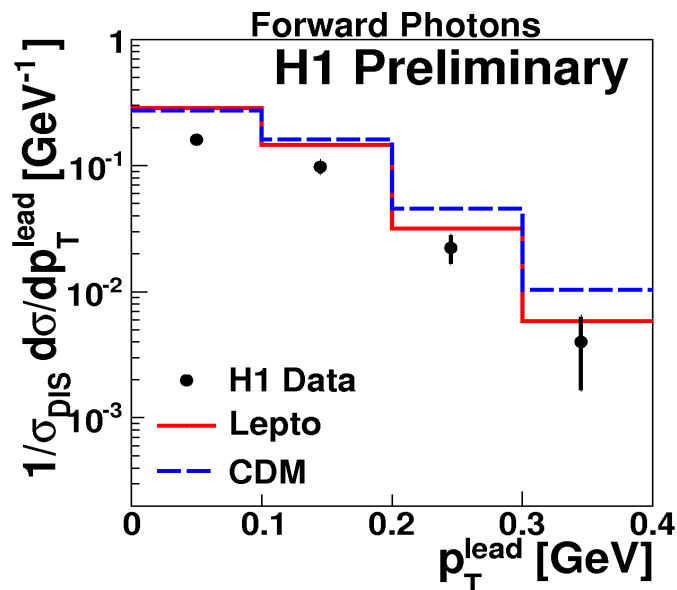
Forward photon production cross sections vs x_L



Photon rate in all tested Monte Carlo models is significantly higher than in data.

LEPTO describes the shape reasonably well.

p_T^2 spectrum shape is well described by SIBYLL and EPOS models.



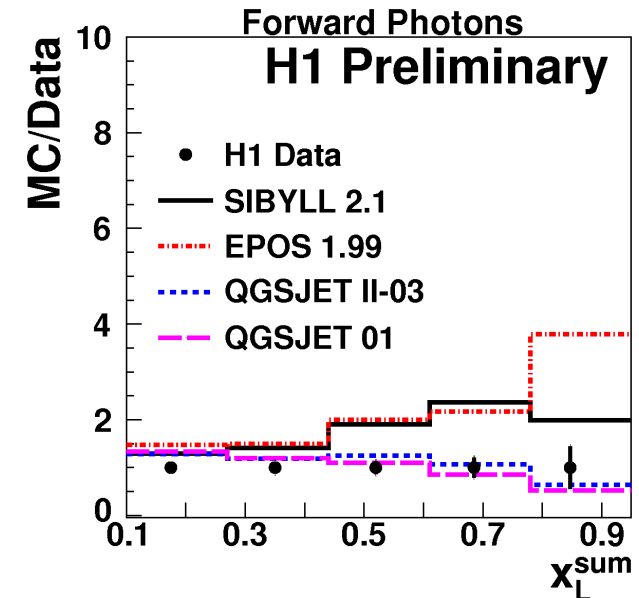
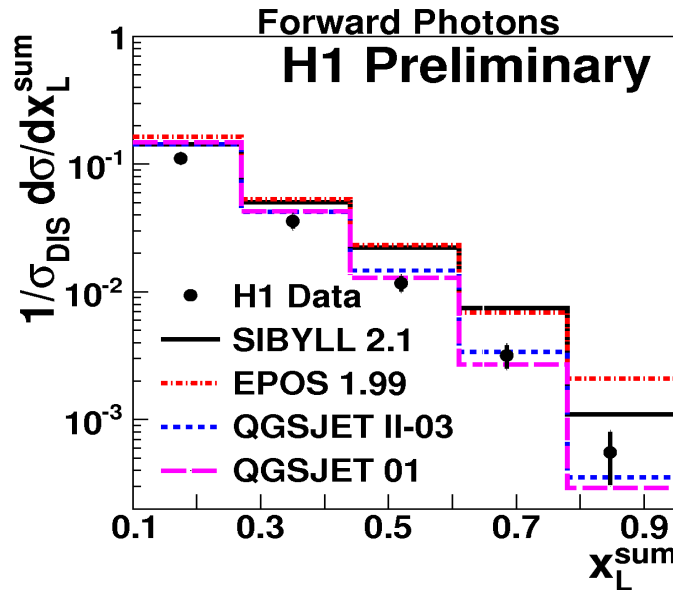
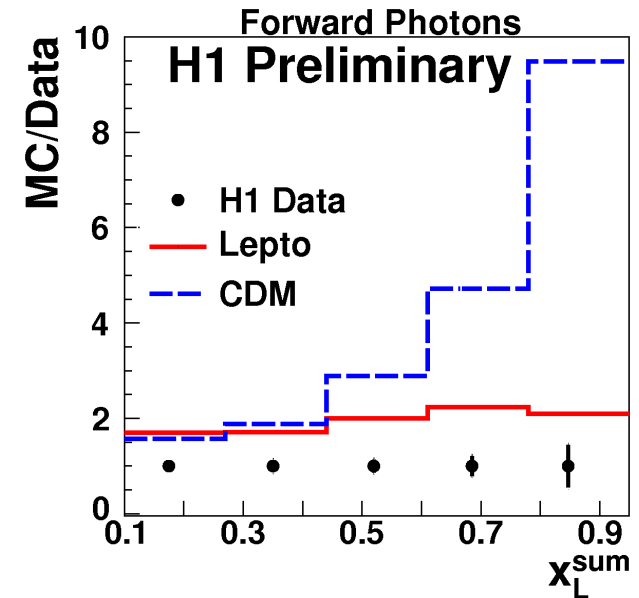
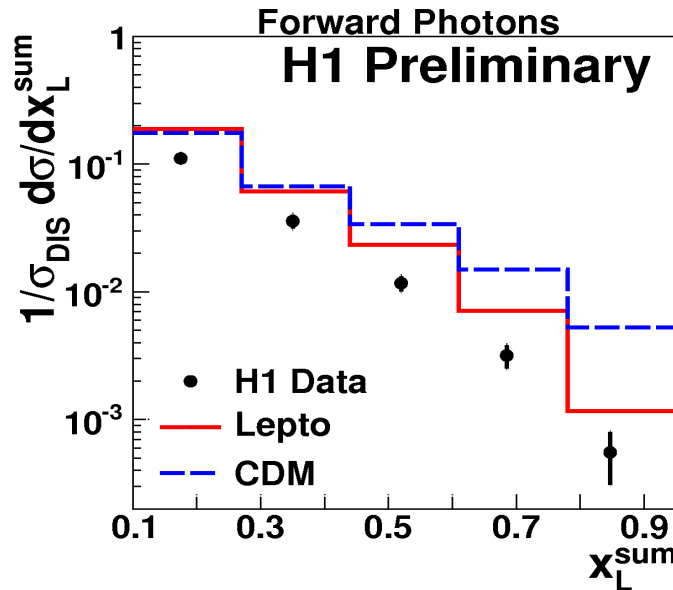
$$x_L^{\text{sum}} = \sum E_i / E_{p\text{-beam}}$$

of γ with $\eta > 7.9$

Photon rate in all tested Monte Carlo models is significantly higher than in data.

LEPTO describes the shape reasonably well. At higher x_L CDM to data ratio is even worse.

For energy sum QGSJET models describe data shape better than SIBYLL and EPOS.



Summary

- ◆ Measurement of double-differential cross section vs x_L and p_T^2 of leading neutron production is presented in DIS.
- ◆ Fragmentation MC-models without meson exchange do not describe the data, addition of model with pion exchange mechanism allows a better description of the data.
- ◆ Pion flux can be further constrained using the measurement.

- ◆ Precise measurements of leading neutron x_L and p_T^2 presented in γp with dijets.
- ◆ Reintroducing hard scale in γp with high E_T jets: absorption effect not prominent.
- ◆ Unlike Monte Carlo predictions x_γ distributions ratio (dijet+LN)/(dijet) is not flat in data.
- ◆ Leading neutron production in γp rate is suppressed at $x_L < 0.85$.

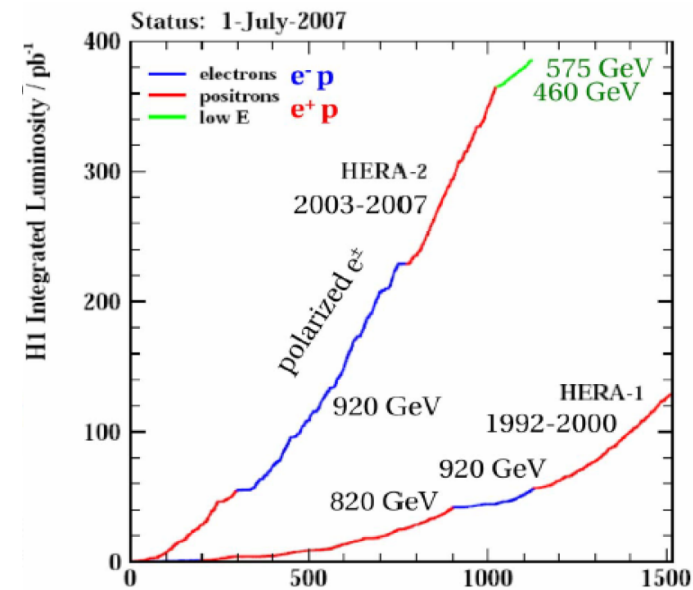
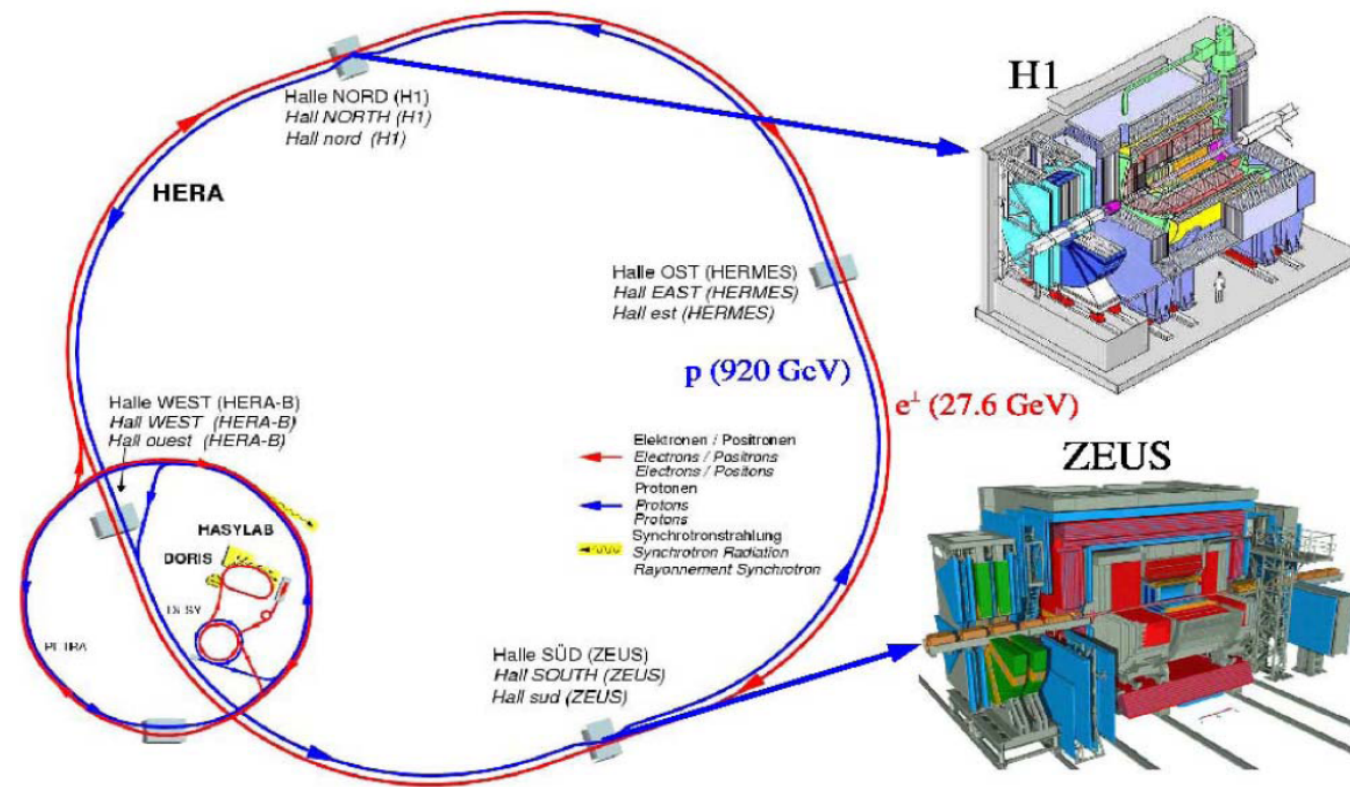
- ◆ First measurement of $1/\sigma_{DIS}$ normalized cross section of Forward Photons production in DIS.
- ◆ Measurements show sensitivity to proton fragmentation MC models.
- ◆ Photon rate in Monte Carlo models is significantly higher than in data.
- ◆ LEPTO describes the shape reasonably well, CDM predicts too many photons at high energies
- ◆ QGSJET models provide reasonable description of x_L dependence while EPOS and SIBYLL provide similar level of description of p_T dependence.
- ◆ None of the models describes data either in normalization or in shape.

Backup

HERA

The world's only electron/positron-proton collider at DESY, Hamburg.

$E_e = 27.6 \text{ GeV}$, $E_p = 920 \text{ GeV}$ (also 820, 460 and 575 GeV). \sqrt{s} up to 320 GeV.



HERA-1: 1992 - 2000
HERA-2: 2003 - 2007

Two colliding experiments: H1 and ZEUS

Total lumi: 0.5 fb^{-1} per experiment

$$e+p \rightarrow e+n+X$$

$$\text{Fit by: } \frac{1}{\sigma_{inc}} \frac{d\sigma_{LN}}{dp_T^2 dx_L} = a(x_L) \cdot e^{-b(x_L) p_T^2}$$

Slopes $b(x_L)$ as a function of x_L
compared to several selected pion
flux parameterizations:

KPP:

B.Kopeliovich, B.Povh, I.Potsahnikova

Bishari:

M.Bishari

Holtmann:

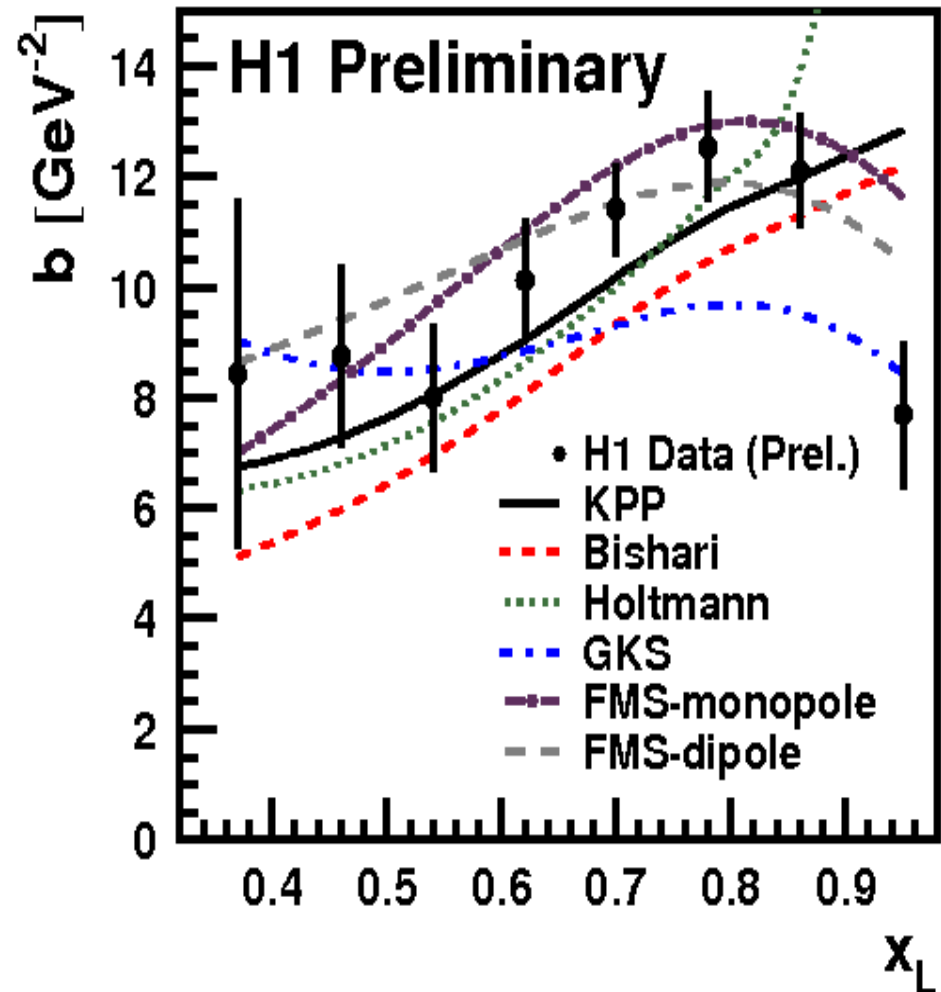
H.Holtmann

GKS:

K.J.Golec-Biernat, J.Kwiecinski, A.Szczurek

FMS:

L.L.Frankfurt, L.Mankiewicz, M.I.Strikman



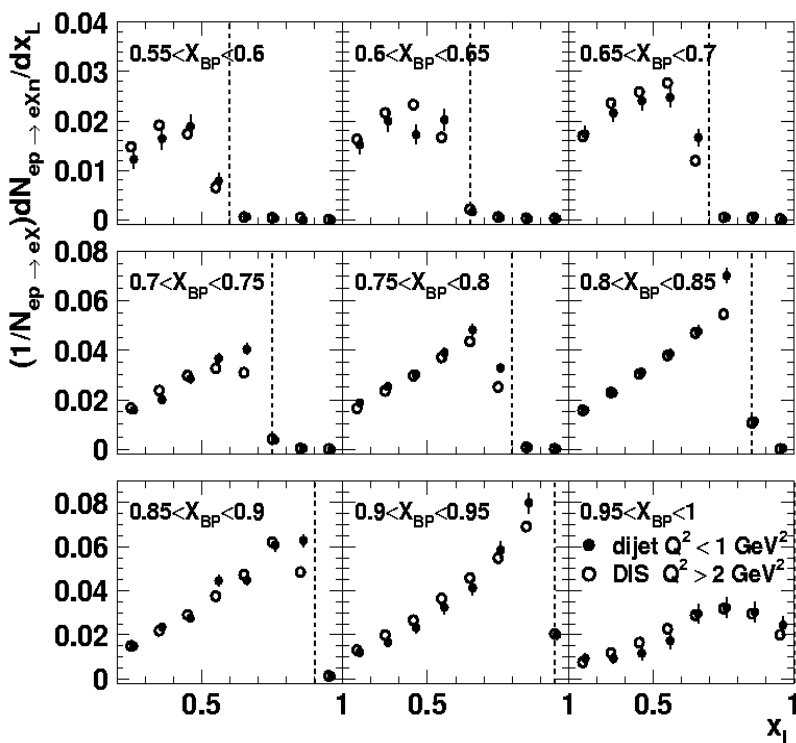
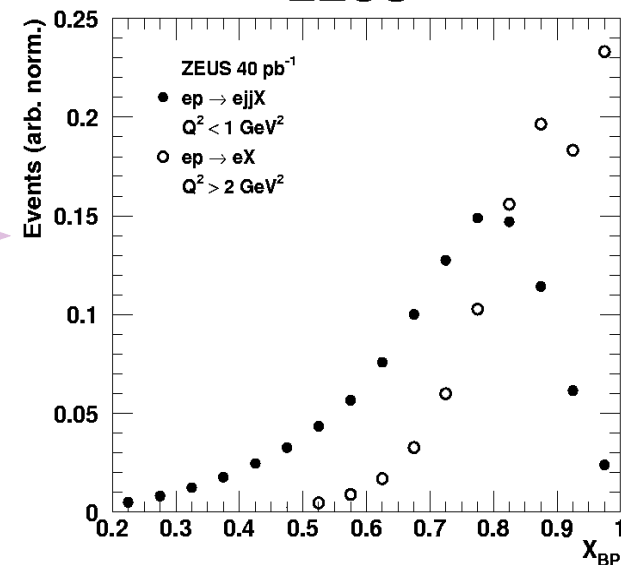
Consider X_{BP} = fraction of p-energy available for LN production

$$x_L < X_{BP} = 1 - (E + P_Z)/(2E_p)$$

X_{BP} dist. is different in DIS and dijet γp :
much less energy available in
dijet γp for LN production

ZEUS

ZEUS



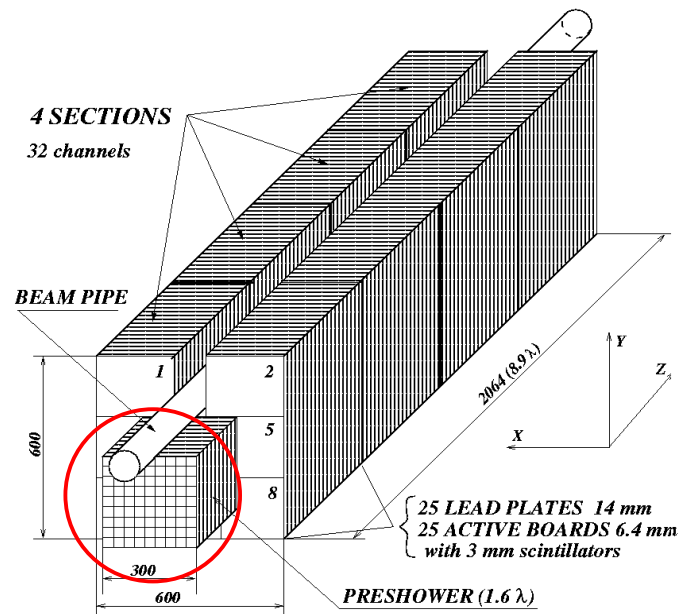
Reweight DIS LN x_L dist. to match the X_{BP} dist. in dijet γp

suppression at high x_L dist. mostly gone
large suppression at low x_L seen in γp
without jets not there

For fixed X_{BP} , same LN rate and x_L spectrum

Differences in the x_L spectra
due to kinematic suppression.

Preshower detector



- $26 \times 26 \times 40 \text{ cm}^3$
- $60 X_0, 1.6 \lambda_i$
- readout 9 X and 9 Y projection strips

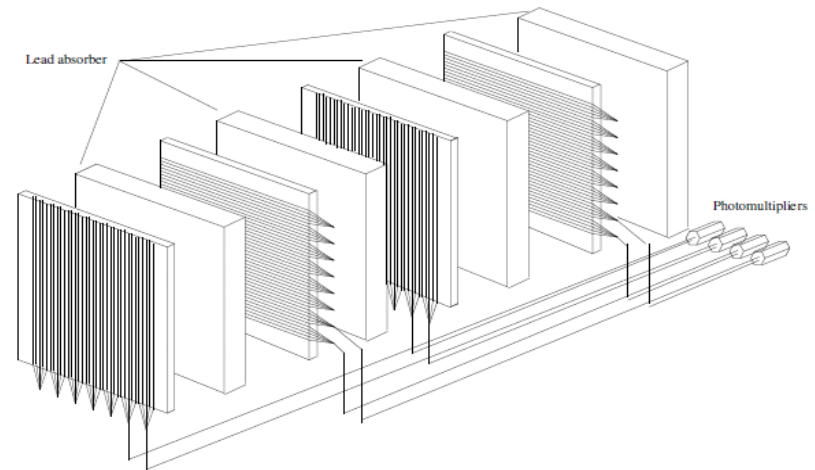
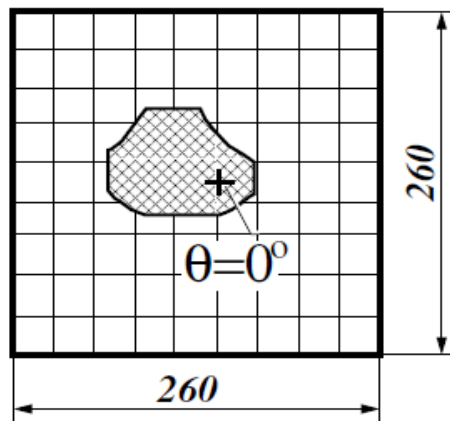


Figure 3: The scheme of light collection for the Preshower.



Segmentation in depth (Z-axis) allows reliable discrimination between electromagnetic and hadronic showers, that is between photons and neutrons.

At low energy ($x_L < 0.1$) neutrons can be misidentified as photons \Rightarrow measure cross sections for $x_L > 0.1$